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# Deformed shape invariance and exactly solvable Hamiltonians with position-dependent effective mass 

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#### Abstract

Known shape-invariant potentials for the constant-mass Schrödinger equation are taken as effective potentials in a position-dependent effective mass (PDEM) one. The corresponding shape-invariance condition turns out to be deformed. Its solvability imposes the form of both the deformed superpotential and the PDEM. A lot of new exactly solvable potentials associated with a PDEM background are generated in this way. A novel and important condition restricting the existence of bound states whenever the PDEM vanishes at an end point of the interval is identified. In some cases, the bound-state spectrum results from a smooth deformation of that of the conventional shape-invariant potential used in the construction. In others, one observes a generation or suppression of bound states, depending on the mass-parameter values. The corresponding wavefunctions are given in terms of some deformed classical orthogonal polynomials.


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## 1. Introduction

There has been a growing interest in studying position-dependent effective mass (PDEM) quantum Hamiltonians due to their relevance in describing the dynamics of electrons in many condensed-matter systems, such as compositionally graded crystals [1], quantum dots [2] and liquid crystals [3]. The PDEM concept has been considered in the energy-density functional approach to the quantum many-body problem in the context of nonlocal terms
of the accompanying potential and applied to nuclei [4], quantum liquids [5] and metal clusters [6], for instance. Some other theoretical advances include the derivation of the underlying electron Hamiltonian from instantaneous Galilean invariance [7] and the calculation of Green's function for step and rectangular-barrier potentials and masses [8] by implementing path-integral techniques [9].

Many recent developments have aimed at deriving exact solutions of the PDEM Schrödinger equation (SE) [10-20]. They have been achieved by extending some well-known methods used to generate exactly solvable (ES), quasi-ES or conditionally ES potentials. Such methods include point canonical transformations [21], Lie algebraic methods [22], as well as supersymmetric quantum mechanical (SUSYQM) and shape-invariance (SI) techniques [23, 24].

In a recent paper [18], Quesne and Tkachuk have pointed out certain intimate connections between the PDEM SE and the constant-mass SE based on deformed canonical commutation relations (see also [25] for a treatment on the classical aspect). Their study exploits the existence of a specific relation between the PDEM and the deforming function appearing in the generalized canonical commutation relations. As a consequence of this relation, the potential in the deformed SE may be considered as the effective potential in the PDEM one, taking into account the interplay of the initial potential and the ambiguity-parameterdependent contribution of the kinetic energy term coming from the momentum and massoperator noncommutativity.

In the approach of [18], solving a PDEM SE for a specific choice of the mass function amounts to considering some deformed SI condition in a SUSYQM framework. This relates the PDEM formalism to an important branch of SUSYQM, whose development dates back to that of quantum groups and $q$-algebras and which has produced a lot of interesting results (see, e.g., [26-29]).

The procedure proposed in [18] has been illustrated by considering the case of the threedimensional Coulomb problem bound-state energy spectrum. This example has revealed two interesting features. First, the ambiguity parameters have been shown to essentially lead to reparametrizing the Coulomb potential without changing its shape. Second, a drastic effect of the mass environment on the energy spectrum has been uncovered in the sense that the infinite bound-state spectrum of the constant-mass case is converted into a finite one.

Both of these results strongly contrast with those of most constructions of solvable PDEM SE's, where the potential gets mass deformed in a rather complicated way while the spectrum remains the same as in the constant-mass case. One notable exception to this general observation comes from a recent analysis of the free-particle problem, where the presence of a suitable mass environment generates an infinite number of bound states [19].

In this paper, our primary concern is to extend the procedure of [18] to those onedimensional potentials that are SI under parameter translation [24]. We actually plan to show that under some suitable assumptions on the corresponding superpotential, one may find a PDEM or, equivalently, a deforming function, for which the deformed SI condition remains solvable, thereby leading to exact results for the bound-state spectrum and the corresponding wavefunctions of the associated SE's, provided the latter satisfy some appropriate conditions. Our secondary purposes consist in studying the interplay of the two contributions to the effective potential and the generation of the corresponding ES PDEM potential, as well as in determining whether the associated mass function has a dramatic or only smooth effect on the bound-state spectrum.

In section 2, the general procedure for solving PDEM SE's through the use of a deformed SI condition is reviewed. In section 3, various classes of superpotentials are identified. The method is then illustrated in section 4 by considering some simple examples. The general
results, listed in the appendix, are commented in section 5. Finally, section 6 contains the conclusion.

## 2. General procedure

One of the well-known problems of the PDEM SE consists of the momentum and massoperator noncommutativity and the resultant ordering ambiguity in the kinetic energy term (see, e.g., [7, 30-32]). To cope with this difficulty, it is advantageous to use the von Roos general two-parameter form of the effective-mass kinetic energy operator [33], which has an inbuilt Hermiticity and contains other plausible forms as special cases.

In units wherein $\hbar=2 m_{0}=1$, we may therefore write the PDEM SE as

$$
\begin{align*}
& {\left[-\frac{1}{2}\left(M^{\xi^{\prime}}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} M^{\eta^{\prime}}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} M^{\zeta^{\prime}}(\boldsymbol{\alpha} ; x)\right.\right.} \\
& \left.\left.\quad+M^{\zeta^{\prime}}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} M^{\eta^{\prime}}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} M^{\xi^{\prime}}(\boldsymbol{\alpha} ; x)\right)+V(\boldsymbol{a} ; x)\right] \psi(x)=E \psi(x) \tag{2.1}
\end{align*}
$$

where $M(\boldsymbol{\alpha} ; x)$ is the dimensionless form of the mass function $m(\boldsymbol{\alpha} ; x)=m_{0} M(\boldsymbol{\alpha} ; x), \boldsymbol{\alpha}$ and $\boldsymbol{a}$ denote two sets of parameters, and the von Roos ambiguity parameters $\xi^{\prime}, \eta^{\prime}, \zeta^{\prime}$ are constrained by the condition $\xi^{\prime}+\eta^{\prime}+\zeta^{\prime}=-1$.

On setting

$$
\begin{equation*}
M(\boldsymbol{\alpha} ; x)=\frac{1}{f^{2}(\boldsymbol{\alpha} ; x)} \quad f(\boldsymbol{\alpha} ; x)=1+g(\boldsymbol{\alpha} ; x) \tag{2.2}
\end{equation*}
$$

where $f(\boldsymbol{\alpha} ; x)$ is some positive-definite function and $g(\boldsymbol{\alpha} ; x)=0$ corresponds to the constantmass case, equation (2.1) becomes

$$
\begin{align*}
& {\left[-\frac{1}{2}\left(f^{\xi}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} f^{\eta}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} f^{\zeta}(\boldsymbol{\alpha} ; x)\right.\right.} \\
& \left.\left.\quad+f^{\zeta}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} f^{\eta}(\boldsymbol{\alpha} ; x) \frac{\mathrm{d}}{\mathrm{~d} x} f^{\xi}(\boldsymbol{\alpha} ; x)\right)+V(\boldsymbol{a} ; x)\right] \psi(x)=E \psi(x) \tag{2.3}
\end{align*}
$$

with $\xi+\eta+\zeta=2$. Among those ambiguity parameter choices that have been found useful for describing the motion of electrons in compositionally graded crystals, we may quote those of BenDaniel and Duke (BDD) [34] $(\xi=0, \zeta=0)$, Bastard [35] $(\xi=2, \zeta=0)$, Zhu and Kroemer (ZK) [36] ( $\xi=1, \zeta=1$ ) and Li and $\operatorname{Kuhn}(\mathrm{LK})[37](\xi=0, \zeta=1)$.

We can get rid of the ambiguity parameters $\xi, \eta, \zeta$ (denoted collectively by $\boldsymbol{\xi}$ ) in the kinetic energy term by transferring them to the effective potential energy of the variable-mass system. Thus using the result

$$
\begin{align*}
f^{\xi} \frac{\mathrm{d}}{\mathrm{~d} x} f^{\eta} \frac{\mathrm{d}}{\mathrm{~d} x} f^{\zeta} & +f^{\zeta} \frac{\mathrm{d}}{\mathrm{~d} x} f^{\eta} \frac{\mathrm{d}}{\mathrm{~d} x} f^{\xi}=2 \sqrt{f} \frac{\mathrm{~d}}{\mathrm{~d} x} f \frac{\mathrm{~d}}{\mathrm{~d} x} \sqrt{f}-(1-\xi-\zeta) f f^{\prime \prime} \\
& -2\left(\frac{1}{2}-\xi\right)\left(\frac{1}{2}-\zeta\right) f^{\prime 2} \tag{2.4}
\end{align*}
$$

where a prime denotes derivative with respect to $x$ and the positive definiteness of $f$ is explicitly used, equation (2.3) acquires the form
$H \psi(x) \equiv\left[-\left(\sqrt{f(\boldsymbol{\alpha} ; x)} \frac{\mathrm{d}}{\mathrm{d} x} \sqrt{f(\boldsymbol{\alpha} ; x)}\right)^{2}+V_{\mathrm{eff}}(\boldsymbol{b} ; x)\right] \psi(x)=E \psi(x)$
in which the effective potential

$$
\begin{equation*}
V_{\mathrm{eff}}(\boldsymbol{b} ; x)=V(\boldsymbol{a} ; x)+\tilde{V}(\boldsymbol{\alpha}, \boldsymbol{\xi} ; x) \tag{2.6}
\end{equation*}
$$

contains an additional mass- and ambiguity-parameter-depending term

$$
\begin{equation*}
\tilde{V}(\boldsymbol{\alpha}, \boldsymbol{\xi} ; x)=\rho f(\boldsymbol{\alpha} ; x) f^{\prime \prime}(\boldsymbol{\alpha} ; x)+\sigma f^{\prime 2}(\boldsymbol{\alpha} ; x) \tag{2.7}
\end{equation*}
$$

In (2.5) and (2.6), the parameters $\boldsymbol{b}$ depend on the whole set of parameters $\boldsymbol{a}, \boldsymbol{\alpha}$ and $\boldsymbol{\xi}$, while in (2.7) we have denoted by $\rho$ and $\sigma$ the following two ambiguity-parameter combinations:

$$
\begin{equation*}
\rho=\frac{1}{2}(1-\xi-\zeta) \quad \sigma=\left(\frac{1}{2}-\xi\right)\left(\frac{1}{2}-\zeta\right) \tag{2.8}
\end{equation*}
$$

For the special ambiguity-parameter choices referred to hereabove, they take the values $\rho=\frac{1}{2}, \sigma=\frac{1}{4}$ (BDD), $\rho=-\frac{1}{2}, \sigma=-\frac{3}{4}$ (Bastard), $\rho=-\frac{1}{2}, \sigma=\frac{1}{4}$ (ZK), or $\rho=0, \sigma=-\frac{1}{4}$ (LK).

The PDEM SE (2.5) may now be reinterpreted as a deformed SE

$$
\begin{equation*}
H \psi(x)=\left[\pi^{2}+V_{\mathrm{eff}}(\boldsymbol{b} ; x)\right] \psi(x)=E \psi(x) \tag{2.9}
\end{equation*}
$$

corresponding to the replacement of the momentum operator $p=-\mathrm{i} \frac{\mathrm{d}}{\mathrm{d} x}$ by some deformed one

$$
\begin{equation*}
\pi \equiv \sqrt{f(\boldsymbol{\alpha} ; x)} p \sqrt{f(\boldsymbol{\alpha} ; x)}=-\mathrm{i} \sqrt{f(\boldsymbol{\alpha} ; x)} \frac{\mathrm{d}}{\mathrm{~d} x} \sqrt{f(\boldsymbol{\alpha} ; x)} . \tag{2.10}
\end{equation*}
$$

With this substitution, the conventional commutation relation $[x, p]=\mathrm{i}$ is changed into

$$
\begin{equation*}
[x, \pi]=\mathrm{i} f(\boldsymbol{\alpha} ; x) \tag{2.11}
\end{equation*}
$$

where $f(\boldsymbol{\alpha} ; x)$ acts as a deforming function.
In this paper, we plan to take for $V_{\text {eff }}(\boldsymbol{b} ; x)$ some known SI potential. This means that the initial potential in the PDEM SE (2.3) will then be determined by inverting (2.6) as

$$
\begin{equation*}
V(\boldsymbol{a} ; x)=V_{\mathrm{eff}}(\boldsymbol{b} ; x)-\tilde{V}(\boldsymbol{\alpha}, \boldsymbol{\xi} ; x) \tag{2.12}
\end{equation*}
$$

where the parameters $\boldsymbol{a}$ now depend on the SI potential parameters $\boldsymbol{b}$ and on $\boldsymbol{\alpha}, \boldsymbol{\xi}$.
To solve equation (2.9) (and therefore (2.3)), we will show that for some appropriately chosen deforming function $f(\boldsymbol{\alpha} ; x), H$ may be considered as the first member $H_{0}=H$ of a hierarchy of Hamiltonians

$$
\begin{equation*}
H_{i}=A^{+}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right) A^{-}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right)+\sum_{j=0}^{i} \epsilon_{j} \quad i=0,1,2, \ldots \tag{2.13}
\end{equation*}
$$

where the first-order operators

$$
\begin{equation*}
A^{ \pm}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right)=\mp \sqrt{f(\boldsymbol{\alpha} ; x)} \frac{\mathrm{d}}{\mathrm{~d} x} \sqrt{f(\boldsymbol{\alpha} ; x)}+W\left(\boldsymbol{\lambda}_{i} ; x\right) \tag{2.14}
\end{equation*}
$$

satisfy a deformed SI condition
$A^{-}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right) A^{+}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right)=A^{+}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i+1}\right) A^{-}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i+1}\right)+\epsilon_{i+1} \quad i=0,1,2, \ldots$
and $\epsilon_{i}, i=0,1,2, \ldots$, are some constants. It follows from equation (2.15) that we can rewrite $H_{i+1}$ as

$$
\begin{equation*}
H_{i+1}=A^{-}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right) A^{+}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right)+\sum_{j=0}^{i} \epsilon_{j} \quad i=0,1,2, \ldots \tag{2.16}
\end{equation*}
$$

so that Hamiltonians (2.13) fulfil intertwining relations

$$
\begin{equation*}
H_{i} A^{+}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right)=A^{+}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right) H_{i+1} \quad A^{-}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right) H_{i}=H_{i+1} A^{-}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{i}\right) \tag{2.17}
\end{equation*}
$$

similar to those of the undeformed case.

Solving equation (2.15) means that it is possible to find a superpotential $W(\lambda ; x)$, a deforming function $f(\boldsymbol{\alpha} ; x)$ and some constants $\boldsymbol{\lambda}_{i}, \epsilon_{i}, i=0,1,2, \ldots$, with $\boldsymbol{\lambda}_{0}=\boldsymbol{\lambda}$, such that

$$
\begin{equation*}
V_{\mathrm{eff}}(\boldsymbol{b} ; x)=W^{2}(\boldsymbol{\lambda} ; x)-f(\boldsymbol{\alpha} ; x) W^{\prime}(\boldsymbol{\lambda} ; x)+\epsilon_{0} \tag{2.18}
\end{equation*}
$$

and

$$
\begin{align*}
W^{2}\left(\boldsymbol{\lambda}_{i} ; x\right)+ & f(\boldsymbol{\alpha} ; x) W^{\prime}\left(\boldsymbol{\lambda}_{i} ; x\right) \\
& =W^{2}\left(\boldsymbol{\lambda}_{i+1} ; x\right)-f(\boldsymbol{\alpha} ; x) W^{\prime}\left(\boldsymbol{\lambda}_{i+1} ; x\right)+\epsilon_{i+1} \quad i=0,1,2, \ldots \tag{2.19}
\end{align*}
$$

As a consequence, the (deformed) SUSY partner $H_{1}$ of $H$ will be characterized by a potential

$$
\begin{equation*}
V_{\mathrm{eff}, 1}(\boldsymbol{b}, \boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=V_{\mathrm{eff}}(\boldsymbol{b} ; x)+2 f(\boldsymbol{\alpha} ; x) W^{\prime}(\boldsymbol{\lambda} ; x) \tag{2.20}
\end{equation*}
$$

To find a solution to equations (2.18) and (2.19), we shall be guided by our knowledge of the superpotential $W$ in the undeformed case $(f=1$ or $g=0)$ [24], where the parameters $\boldsymbol{\lambda}$ are entirely determined by the potential parameters $\boldsymbol{b}$. Our strategy will consist in (i) assuming that the deformation does not affect the form of $W$ but only brings about a change in its parameters $\boldsymbol{\lambda}$ (which will now also depend on $\boldsymbol{\alpha}$ ), and (ii) choosing $g(\boldsymbol{\alpha} ; x)$ in such a way that in (2.18) and (2.19) the function $g(\boldsymbol{\alpha} ; x) W^{\prime}(\boldsymbol{\lambda} ; x)$ contains the same kind of terms as those already present in the undeformed case, i.e., $W^{2}(\boldsymbol{\lambda} ; x)$ and $W^{\prime}(\boldsymbol{\lambda} ; x)$. In section 3, we shall put this recipe into practice for general classes of superpotentials and determine the accompanying deforming function $f(\boldsymbol{\alpha} ; x)$, from which the corresponding PDEM can then be obtained through equation (2.2).

It is worth noting that although on solving equation (2.18), $\boldsymbol{\lambda}$ will become a known function of $\boldsymbol{b}$ and $\boldsymbol{\alpha}$, it will often prove convenient to keep it as a (redundant) argument in operators, energies and wavefunctions.

Having found a solution to equations (2.18) and (2.19), we can determine the bound-state energy spectrum and corresponding wavefunctions of $H$ by an extension of the conventional SUSYQM and SI procedure [23, 24]. Thus the energy eigenvalues are given by

$$
\begin{equation*}
E_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda})=\sum_{i=0}^{n} \epsilon_{i} \tag{2.21}
\end{equation*}
$$

while the ground- and excited-state wavefunctions are obtained by solving the first-order differential equation

$$
\begin{equation*}
A^{-}(\boldsymbol{\alpha}, \boldsymbol{\lambda}) \psi_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=0 \tag{2.22}
\end{equation*}
$$

and the recursion relation

$$
\begin{equation*}
\psi_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=\left[E_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda})-E_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda})\right]^{-1 / 2} A^{+}(\boldsymbol{\alpha}, \boldsymbol{\lambda}) \psi_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; x\right) \tag{2.23}
\end{equation*}
$$

respectively.
Equations (2.21), (2.22) and (2.23) only provide formal solutions to equation (2.5) or (2.9). To be physically acceptable, the bound-state wavefunctions should indeed satisfy two conditions:
(i) As in conventional quantum mechanics, they should be square integrable on the (finite or infinite) interval of definition of $V_{\text {eff }}(\boldsymbol{b} ; \boldsymbol{x})$, i.e.,

$$
\begin{equation*}
\int_{x_{1}}^{x_{2}} \mathrm{~d} x\left|\psi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)\right|^{2}<\infty \tag{2.24}
\end{equation*}
$$

(ii) Furthermore, they should ensure the Hermiticity of $H$. For such a purpose, it is enough to impose that the deformed momentum operator $\pi$, defined in (2.10), be Hermitian. This amounts to the condition

$$
\begin{align*}
\int_{x_{1}}^{x_{2}} \mathrm{~d} x \psi^{*}(x) & \sqrt{f(\boldsymbol{\alpha} ; x)}\left(-\mathrm{i} \frac{\mathrm{~d}}{\mathrm{~d} x}\right) \sqrt{f(\boldsymbol{\alpha} ; x)} \phi(x) \\
& =\left[\int_{x_{1}}^{x_{2}} \mathrm{~d} x \phi^{*}(x) \sqrt{f(\boldsymbol{\alpha} ; x)}\left(-\mathrm{i} \frac{\mathrm{~d}}{\mathrm{~d} x}\right) \sqrt{f(\boldsymbol{\alpha} ; x)} \psi(x)\right]^{*} \tag{2.25}
\end{align*}
$$

for any $\psi(x), \phi(x) \in L^{2}\left(x_{1}, x_{2}\right)$. Integrating the left-hand side of (2.25) by parts leads to

$$
\begin{align*}
\int_{x_{1}}^{x_{2}} \mathrm{~d} x \psi^{*}(x) & \sqrt{f(\boldsymbol{\alpha} ; x)}\left(-\mathrm{i} \frac{\mathrm{~d}}{\mathrm{~d} x}\right) \sqrt{f(\boldsymbol{\alpha} ; x)} \phi(x)=-\left.\mathrm{i} \psi^{*}(x) \phi(x) f(\boldsymbol{\alpha} ; x)\right|_{x_{1}} ^{x_{2}} \\
& +\int_{x_{1}}^{x_{2}} \mathrm{~d} x \phi(x) \sqrt{f(\boldsymbol{\alpha} ; x)}\left(\mathrm{i} \frac{\mathrm{~d}}{\mathrm{~d} x}\right) \sqrt{f(\boldsymbol{\alpha} ; x)} \psi^{*}(x) \tag{2.26}
\end{align*}
$$

Comparison with the right-hand side of (2.25) then provides us with the condition $\psi^{*}(x) \phi(x) f(\alpha ; x) \rightarrow 0$ for $x \rightarrow x_{1}$ and $x \rightarrow x_{2}$. This shows that one has to place the restriction

$$
\begin{equation*}
\left|\psi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)\right|^{2} f(\boldsymbol{\alpha} ; x) \rightarrow 0 \quad \text { for } \quad x \rightarrow x_{1} \quad \text { and } \quad x \rightarrow x_{2} \tag{2.27}
\end{equation*}
$$

on the allowed bound-state wavefunctions. This condition will be effective whenever $f(\boldsymbol{\alpha} ; x)$ does not go to some finite constant at the end points of the interval.

The precise range of $n$ values ( $n=0,1, \ldots, n_{\max }$ or $n=0,1,2, \ldots$ ) in equation (2.21) will therefore be determined by the existence of corresponding wavefunctions $\psi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)$ satisfying both equations (2.24) and (2.27). In terms of PDEM (2.2), the latter condition translates into

$$
\begin{equation*}
\frac{\left|\psi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)\right|^{2}}{\sqrt{M(\boldsymbol{\alpha} ; x)}} \rightarrow 0 \quad \text { for } \quad x \rightarrow x_{1} \quad \text { and } \quad x \rightarrow x_{2} \tag{2.28}
\end{equation*}
$$

which should be checked whenever $M(\alpha ; x) \rightarrow 0$ for $x \rightarrow x_{1}$ or $x \rightarrow x_{2}$. It should be stressed that although this condition may be present in any PDEM problem, it has not been noted so far.

On taking (2.14) into account, the solution of equation (2.22) can be formally obtained in terms of $W$ and $f$. It is given by

$$
\begin{equation*}
\psi_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=\frac{N_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda})}{\sqrt{f(\boldsymbol{\alpha} ; x)}} \exp \left(-\int^{x} \frac{W(\boldsymbol{\lambda} ; \tilde{x})}{f(\boldsymbol{\alpha} ; \tilde{x})} \mathrm{d} \tilde{x}\right) \tag{2.29}
\end{equation*}
$$

where $N_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda})$ is some normalization coefficient.
Similarly, the solution of (2.23) can be shown to be

$$
\begin{equation*}
\psi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=\frac{N_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda})}{\sqrt{f(\boldsymbol{\alpha} ; x)}} \varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x) \exp \left(-\int^{x} \frac{W\left(\boldsymbol{\lambda}_{n} ; \tilde{x}\right)}{f(\boldsymbol{\alpha} ; \tilde{x})} \mathrm{d} \tilde{x}\right) \tag{2.30}
\end{equation*}
$$

where $\varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)$ fulfils the equation

$$
\begin{equation*}
\varphi_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=-f(\boldsymbol{\alpha} ; x) \varphi_{n}^{\prime}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; x\right)+\left[W\left(\boldsymbol{\lambda}_{n+1} ; x\right)+W(\boldsymbol{\lambda} ; x)\right] \varphi_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; x\right) \tag{2.31}
\end{equation*}
$$

with $\varphi_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)=1$, and the normalization coefficient $N_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda})$ satisfies the recursion relation

$$
\begin{equation*}
N_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda})=\left[E_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda})-E_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda})\right]^{-1 / 2} N_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1}\right) \tag{2.32}
\end{equation*}
$$

3. Classes of superpotentials and their accompanying deforming function

In this section, we plan to consider several classes of superpotentials, which in the following sections will prove to include all the SI potentials considered in table 4.1 of [24], as well as their special cases. For each class, we shall determine the general form of the accompanying deforming function. From the expressions obtained for $W$ and $f$, we shall then deduce some consequences regarding the ground- and excited-state wavefunction explicit form.

### 3.1. Classes of superpotentials

Let $\phi(x)$ be some parameter-independent function of $x$ and $\boldsymbol{\lambda}$ denote a single parameter (for class 0 ) or a set of two parameters $\lambda, \mu$ (for classes 1,2 and 3 ).

Class 0. The simplest choice of superpotential is a single-parameter one of the type

$$
\begin{equation*}
W(\lambda ; x)=\lambda \phi(x) . \tag{3.1}
\end{equation*}
$$

Conditions (2.18) and (2.19) then contain two parameters to be determined, namely $\lambda, \epsilon_{0}$ and $\lambda_{i+1}, \epsilon_{i+1}$, respectively.

In the undeformed case, apart from a constant term, $\epsilon_{0}$ or $\epsilon_{i+1}$, they include the functions $W^{2}$ and $W^{\prime}$. Since $W^{2}$ is proportional to $\phi^{2}$ and we need two equations to calculate the couple of undetermined parameters, equations (2.18) and (2.19) solvability imposes that $W^{\prime}=\lambda \phi^{\prime}$ be a linear combination of $\phi^{2}$ and a constant. In other words, there must exist some numerical (i.e., parameter-independent) constants $A$ and $B$ such that

$$
\begin{equation*}
\phi^{\prime}(x)=A \phi^{2}(x)+B . \tag{3.2}
\end{equation*}
$$

In the deformed case, equations (2.18) and (2.19) contain in addition a term $W^{\prime} g=\lambda \phi^{\prime} g$. If this term has the same form as the remaining ones, it will not spoil the equations solvability. This amounts to assuming that there exist two $\boldsymbol{\alpha}$-dependent constants $A^{\prime}(\boldsymbol{\alpha})$ and $B^{\prime}(\boldsymbol{\alpha})$ such that

$$
\begin{equation*}
\phi^{\prime}(x) g(\boldsymbol{\alpha} ; x)=A^{\prime}(\boldsymbol{\alpha}) \phi^{2}(x)+B^{\prime}(\boldsymbol{\alpha}) \tag{3.3}
\end{equation*}
$$

On combining (3.3) with (3.2), we get

$$
\begin{equation*}
g(\boldsymbol{\alpha} ; x)=\frac{A^{\prime}(\boldsymbol{\alpha}) \phi^{2}(x)+B^{\prime}(\boldsymbol{\alpha})}{A \phi^{2}(x)+B} \tag{3.4}
\end{equation*}
$$

which provides us with the general form of the deforming function $f(\boldsymbol{\alpha} ; x)$ for class 0 superpotentials.

Class 1. The most straightforward generalization of (3.1) consists in adding some nonvanishing parameter $\mu$ :

$$
\begin{equation*}
W(\lambda ; x)=\lambda \phi(x)+\mu . \tag{3.5}
\end{equation*}
$$

Equations (2.18) and (2.19) now contain three parameters ( $\lambda, \mu, \epsilon_{0}$ or $\lambda_{i+1}, \mu_{i+1}, \epsilon_{i+1}$ ) to be determined, but as a counterpart $W^{2}$ is also made of three terms proportional to $\phi^{2}, \phi$ and a constant, respectively. We shall then get three equations to govern the parameter values both in the undeformed and deformed cases provided

$$
\begin{equation*}
\phi^{\prime}(x)=A \phi^{2}(x)+B \phi(x)+C \tag{3.6}
\end{equation*}
$$

and

$$
\begin{equation*}
g(\boldsymbol{\alpha} ; x)=\frac{A^{\prime}(\boldsymbol{\alpha}) \phi^{2}(x)+B^{\prime}(\boldsymbol{\alpha}) \phi(x)+C^{\prime}(\boldsymbol{\alpha})}{A \phi^{2}(x)+B \phi(x)+C} \tag{3.7}
\end{equation*}
$$

where $A, B, C$ and $A^{\prime}(\boldsymbol{\alpha}), B^{\prime}(\boldsymbol{\alpha}), C^{\prime}(\boldsymbol{\alpha})$ are some numerical and $\boldsymbol{\alpha}$-dependent constants, respectively.

Comparison of equations (3.1), (3.2), (3.4) with equations (3.5)-(3.7) shows that class 0 superpotentials may be considered as special cases of class 1 superpotentials, corresponding to the simultaneous vanishing of $\mu, B$ and $B^{\prime}(\boldsymbol{\alpha})$. In the following, we shall therefore include class 0 into class 1 by assuming that for the latter either $\mu \neq 0$ or $\mu=B=B^{\prime}(\boldsymbol{\alpha})=0$.

Class 2. If we define $W$ as

$$
\begin{equation*}
W(\boldsymbol{\lambda} ; x)=\lambda \phi(x)+\frac{\mu}{\phi(x)} \tag{3.8}
\end{equation*}
$$

where $\lambda$ and $\mu$ are both nonvanishing (otherwise we would get back class 0 ), $W^{2}$ again contains three terms proportional to $\phi^{2}, \phi^{-2}$ and a constant, respectively.

A reasoning similar to that carried out for class 1 superpotentials leads to the following expressions for $\phi^{\prime}$ and $g$ :

$$
\begin{align*}
& \phi^{\prime}(x)=A \phi^{2}(x)+B  \tag{3.9}\\
& g(\boldsymbol{\alpha} ; x)=\frac{A^{\prime}(\boldsymbol{\alpha}) \phi^{2}(x)+B^{\prime}(\boldsymbol{\alpha})}{A \phi^{2}(x)+B} \tag{3.10}
\end{align*}
$$

where $A, B, A^{\prime}(\boldsymbol{\alpha})$ and $B^{\prime}(\boldsymbol{\alpha})$ are independent of $x$.

Class 3. On assuming

$$
\begin{equation*}
W(\boldsymbol{\lambda} ; x)=\frac{\lambda \phi(x)+\mu}{\sqrt{A \phi^{2}(x)+B}} \tag{3.11}
\end{equation*}
$$

where $\lambda, \mu$ are nonvanishing and $A, B$ are two numerical nonvanishing constants (otherwise we would get back one of the previous classes), we obtain after a simple calculation

$$
\begin{align*}
& W^{2}(\lambda ; x)=\frac{\lambda^{2} \phi^{2}(x)+2 \lambda \mu \phi(x)+\mu^{2}}{A \phi^{2}(x)+B}  \tag{3.12}\\
& W^{\prime}(\boldsymbol{\lambda} ; x)=\frac{\lambda B-\mu A \phi(x)}{A \phi^{2}(x)+B} \frac{\phi^{\prime}(x)}{\sqrt{A \phi^{2}(x)+B}} . \tag{3.13}
\end{align*}
$$

Equations (2.18) and (2.19) solvability in the undeformed case then implies that

$$
\begin{equation*}
\phi^{\prime}(x)=[C \phi(x)+D] \sqrt{A \phi^{2}(x)+B} \tag{3.14}
\end{equation*}
$$

in terms of two additional numerical constants $C$ and $D$. In the deformed case, the supplementary term must therefore be given by

$$
\begin{equation*}
\phi^{\prime}(x) g(\boldsymbol{\alpha} ; x)=\left[C^{\prime}(\boldsymbol{\alpha}) \phi(x)+D^{\prime}(\boldsymbol{\alpha})\right] \sqrt{A \phi^{2}(x)+B} \tag{3.15}
\end{equation*}
$$

where $C^{\prime}(\boldsymbol{\alpha})$ and $D^{\prime}(\boldsymbol{\alpha})$ depend on the deforming parameters $\boldsymbol{\alpha}$. Hence we obtain

$$
\begin{equation*}
g(\boldsymbol{\alpha} ; x)=\frac{C^{\prime}(\boldsymbol{\alpha}) \phi(x)+D^{\prime}(\boldsymbol{\alpha})}{C \phi(x)+D} \tag{3.16}
\end{equation*}
$$

### 3.2. Corresponding wavefunctions

In equation (2.29), the ground-state wavefunction $\psi_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)$ of $H$ is formally given in terms of the integral of the function $W(\boldsymbol{\lambda} ; \tilde{x}) / f(\boldsymbol{\alpha} ; \tilde{x})$. On taking the explicit forms of $W$ and $f$ obtained in section 3.1 into account, it is straightforward to obtain

$$
\begin{array}{rlr}
\int^{x} \frac{W(\boldsymbol{\lambda} ; \tilde{x})}{f(\boldsymbol{\alpha} ; \tilde{x})} \mathrm{d} \tilde{x} & =\int^{\phi(x)} \frac{\lambda \tilde{\phi}+\mu}{\left[A+A^{\prime}(\boldsymbol{\alpha})\right] \tilde{\phi}^{2}+\left[B+B^{\prime}(\boldsymbol{\alpha})\right] \tilde{\phi}+C+C^{\prime}(\boldsymbol{\alpha})} \mathrm{d} \tilde{\phi} & \text { for class } 1 \\
& =\int^{\phi(x)} \frac{\lambda \tilde{\phi}^{2}+\mu}{\tilde{\phi}\left\{\left[A+A^{\prime}(\boldsymbol{\alpha})\right] \tilde{\phi}^{2}+B+B^{\prime}(\boldsymbol{\alpha})\right\}} \mathrm{d} \tilde{\phi} & \text { for class } 2 \\
& =\int^{\phi(x)} \frac{\lambda \tilde{\phi}+\mu}{\left(A \tilde{\phi}^{2}+B\right)\left\{\left[C+C^{\prime}(\boldsymbol{\alpha})\right] \tilde{\phi}+D+D^{\prime}(\boldsymbol{\alpha})\right\}} \mathrm{d} \tilde{\phi} & \text { for class 3 } \tag{3.17}
\end{array}
$$

thus showing that in all three cases the integral can be explicitly carried out by simple integration techniques as in the undeformed case.

Furthermore, it is possible to write the functions $\varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)$, entering the general expression (2.30) of excited-state wavefunctions, in terms of $n$ th-degree polynomials in a new variable $y, P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)$, which fulfil some equation deriving from (2.31). This result generalizes to the deformed case a well-known property according to which SI potential wavefunctions can be expressed in terms of some classical orthogonal polynomials [24].

The precise form of the changes of variable $x \rightarrow y$ and of function $\varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x) \rightarrow$ $P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)$, as well as the relation satisfied by $P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)$, actually depend on the superpotential class as listed herebelow:

- Class 1

$$
\begin{align*}
\varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)= & P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y) \quad y=\phi(x)  \tag{3.18}\\
P_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)= & -\left\{\left[A+A^{\prime}(\boldsymbol{\alpha})\right] y^{2}+\left[B+B^{\prime}(\boldsymbol{\alpha})\right] y+C+C^{\prime}(\boldsymbol{\alpha})\right\} \dot{P}_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right) \\
& +\left[\left(\lambda_{n+1}+\lambda\right) y+\mu_{n+1}+\mu\right] P_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right) \tag{3.19}
\end{align*}
$$

- Class 2

$$
\begin{align*}
\varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)= & y^{-n / 2} P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y) \quad y=\phi^{-2}(x)  \tag{3.20}\\
P_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)= & 2 y\left\{A+A^{\prime}(\boldsymbol{\alpha})+\left[B+B^{\prime}(\boldsymbol{\alpha})\right] y\right\} \dot{P}_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right) \\
& +\left\{\lambda_{n+1}+\lambda-n\left[A+A^{\prime}(\boldsymbol{\alpha})\right]+\left[\mu_{n+1}+\mu-n\left(B+B^{\prime}(\boldsymbol{\alpha})\right)\right] y\right\} \\
& \times P_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right) . \tag{3.21}
\end{align*}
$$

- Class 3

$$
\begin{align*}
\varphi_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; x)= & \left(A y^{2}+B\right)^{-n / 2} P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y) \quad y=\phi(x)  \tag{3.22}\\
P_{n+1}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)= & \left\{\left[C+C^{\prime}(\boldsymbol{\alpha})\right] y+D+D^{\prime}(\boldsymbol{\alpha})\right\} \\
& \times\left[-\left(A y^{2}+B\right) \dot{P}_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right)+n A y P_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right)\right] \\
& +\left[\left(\lambda_{n+1}+\lambda\right) y+\mu_{n+1}+\mu\right] P_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right) . \tag{3.23}
\end{align*}
$$

Here a dot stands for derivative with respect to $y$ and in all cases $P_{0}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y) \equiv 1$. It should be noted that in (3.23), the linear combination $-\left(A y^{2}+B\right) \dot{P}_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right)+n A y P_{n}\left(\boldsymbol{\alpha}, \boldsymbol{\lambda}_{1} ; y\right)$ is actually an $n$ th-degree polynomial in $y$ because its $(n+1)$ th-degree term vanishes identically.

## 4. Some simple examples

The purpose of this section is twofold: we demonstrate by means of some simple examples how our method developed in the previous sections works in practice and then illustrate the effect of the new restriction (2.27) or (2.28) placed by a deformation or PDEM background on an ES potential bound-state spectrum.

### 4.1. Particle in a box and trigonometric Pöschl-Teller potential

Setting $\mu=0$ and $\phi(x)=\tan x$ in (3.5) leads to the following mutually compatible pair of superpotential and function $g(\alpha ; x)$ :
$W(\lambda ; x)=\lambda \tan x \quad g(\alpha ; x)=\alpha \sin ^{2} x \quad-\frac{\pi}{2} \leqslant x \leqslant \frac{\pi}{2} \quad-1<\alpha \neq 0$
where the range of $\alpha$ restricts the deforming function $f(\alpha ; x)$ to be positive definite in the interval $\left[-\frac{\pi}{2}, \frac{\pi}{2}\right]$, as it should be. Note that with $\lambda=A$ the above superpotential yields in conventional quantum mechanics [38] the familiar trigonometric Pöschl-Teller potential

$$
\begin{equation*}
V_{\mathrm{eff}}(A ; x)=A(A-1) \sec ^{2} x \quad A>1 \quad-\frac{\pi}{2} \leqslant x \leqslant \frac{\pi}{2} \tag{4.2}
\end{equation*}
$$

having the associated bound-state energies and wavefunctions characterized by $E_{n}=(A+n)^{2}$ and $\psi_{n}(x)=N_{n}(\cos x)^{A} C_{n}^{(A)}(\sin x), n=0,1,2, \ldots$ [39]. The particle-in-a-box problem being the limiting case of Pöschl-Teller for $A \rightarrow 1$ corresponds to

$$
V_{\text {eff }}(x)= \begin{cases}0 & \text { if } \quad-\frac{\pi}{2}<x<\frac{\pi}{2}  \tag{4.3}\\ \infty & \text { if } \quad x= \pm \frac{\pi}{2}\end{cases}
$$

Let us first consider the particle in a box. Turning to (2.18) and using (4.1), it is straightforward to obtain the solutions $\lambda=1+\alpha$ and $\epsilon_{0}=1+\alpha$. These imply from (2.19) $\lambda_{i}=(i+1)(1+\alpha)$ and $\epsilon_{i}=(2 i+1)(1+\alpha), i=0,1,2, \ldots$. Consequently, equation (2.21) furnishes the energy eigenvalues

$$
\begin{equation*}
E_{n}(\alpha, \lambda)=(1+\alpha)(n+1)^{2} . \tag{4.4}
\end{equation*}
$$

The corresponding wavefunctions are easily obtainable from (2.29) and (2.30), which give the common form

$$
\begin{equation*}
\psi_{n}(\alpha, \lambda ; x)=N_{n}(\alpha, \lambda) \frac{(\cos x)^{n+1}}{\left(1+\alpha \sin ^{2} x\right)^{(n+2) / 2}} P_{n}(\alpha, \lambda ; \tan x) \tag{4.5}
\end{equation*}
$$

where equations (3.17)-(3.19) have been used. In (4.5), $P_{n}(\alpha, \lambda ; y)$ satisfies the equation
$P_{n+1}(\alpha, \lambda ; y)=-\left[1+(1+\alpha) y^{2}\right] \dot{P}_{n}\left(\alpha, \lambda_{1} ; y\right)+(n+3)(1+\alpha) y P_{n}\left(\alpha, \lambda_{1} ; y\right)$
with $\lambda=1+\alpha$ and $\lambda_{1}=2+2 \alpha$. For $-1<\alpha \neq 0$ and any $n=0,1,2, \ldots$, equation (4.5) manifestly represents a square-integrable function in $\left[-\frac{\pi}{2}, \frac{\pi}{2}\right]$. In addition, since $f\left(\alpha, \pm \frac{\pi}{2}\right)=1+\alpha$, condition (2.27) is also automatically satisfied. We conclude that in the presence of deformation (4.1), the particle-in-a-box problem still has an infinite number of bound states making up a quadratic spectrum. As can be checked, for $\alpha \rightarrow 0$, equations (4.4) and (4.5) go over to their standard forms because $P_{n}(\alpha, \lambda ; \tan x) \rightarrow \gamma_{n} \sec ^{n} x C_{n}^{(1)}(\sin x)$.

When translating this property into the PDEM language, we are led to a new ES SE, corresponding to the mass function given in (2.2), (4.1), and to the potential (2.12), for which

$$
\begin{equation*}
\tilde{V}(\alpha, \rho, \sigma ; x)=-(\rho+\sigma) \alpha^{2} \cos ^{2} 2 x+\rho \alpha(2+\alpha) \cos 2 x+\sigma \alpha^{2} \quad-\frac{\pi}{2} \leqslant x \leqslant \frac{\pi}{2} \tag{4.7}
\end{equation*}
$$

It is worth noting that for the LK choice of ambiguity parameters, the latter expression assumes a very simple form, namely

$$
\begin{equation*}
\tilde{V}(\alpha ; x)=-\frac{1}{4} \alpha^{2} \sin ^{2} 2 x \quad-\frac{\pi}{2} \leqslant x \leqslant \frac{\pi}{2} . \tag{4.8}
\end{equation*}
$$

What has been done for the particle-in-a-box problem can be easily extended to the trigonometric Pöschl-Teller potential. We skip the details, which are lengthy but straightforward, and give the final form of the deformed energy levels and the associated wavefunctions
$E_{n}(\alpha, \lambda)=(\lambda+n)^{2}-\alpha\left(\lambda-n^{2}\right)=\left[\frac{1}{2}(\Delta+1)+n\right]^{2}+\alpha n(n+1)-\frac{1}{4} \alpha^{2}$
$\psi_{n}(\alpha, \lambda ; x)=N_{n}(\alpha, \lambda)(\cos x)^{\frac{\lambda}{1+\alpha}+n}\left(1+\alpha \sin ^{2} x\right)^{-\frac{1}{2}\left(\frac{\lambda}{1+\alpha}+n+1\right)} P_{n}(\alpha, \lambda ; \tan x)$
where $P_{n}(\alpha, \lambda ; y)$ satisfies the equation
$P_{n+1}(\alpha, \lambda ; y)=-\left[1+(1+\alpha) y^{2}\right] \dot{P}_{n}\left(\alpha, \lambda_{1} ; y\right)+[2 \lambda+(n+1)(1+\alpha)] y P_{n}\left(\alpha, \lambda_{1} ; y\right)$.
All functions $\psi_{n}(\alpha, \lambda ; x), n=0,1,2, \ldots$, satisfy both conditions (2.24) and (2.27) again. In the limit $\alpha \rightarrow 0$, the standard results of the Pöschl-Teller are recovered.

### 4.2. Free particle and hyperbolic Pöschl-Teller potential

It is worth comparing what happens for the trigonometric superpotential (4.1) with the case of its hyperbolic counterpart

$$
\begin{equation*}
W(\lambda ; x)=\lambda \tanh x \quad g(\alpha ; x)=\alpha \sinh ^{2} x \quad 0<\alpha<1 \tag{4.12}
\end{equation*}
$$

where the range of $\alpha$ provides us with a positive-definite deforming function $f(\alpha ; x)$ again.
In conventional SUSYQM, the above superpotential has been considered in connection with the attractive or repulsive $\operatorname{sech}^{2} x$ potential (often referred to as the hyperbolic PöschlTeller potential or barrier) [40], as well as with their limiting case, namely the free-particle problem [40, 41]. Let us therefore consider

$$
\begin{equation*}
V_{\mathrm{eff}}(A ; x)=-A(A+1) \operatorname{sech}^{2} x \quad A>0 \tag{4.13}
\end{equation*}
$$

corresponding to the hyperbolic Pöschl-Teller potential (and giving the free-particle problem for $A \rightarrow 0$ ). In the undeformed case, it is known to support a finite number $n_{\max }+$ $1\left(A-1 \leqslant n_{\max }<A\right)$ of bound states, whose energies and wavefunctions are given by $E_{n}=-(A-n)^{2}$ and $\psi_{n}(x)=N_{n}(\operatorname{sech} x)^{A-n} C_{n}^{\left(A-n+\frac{1}{2}\right)}(\tanh x)$, where $n=0,1, \ldots, n_{\max }$ [42]. Such results can be derived by SUSYQM and SI techniques on using the superpotential (4.12) with $\lambda=A$ and the factorization energy $\epsilon_{0}=-A^{2}$ [40].

By proceeding as in section 4.1, in the deformed case we obtain

$$
\begin{equation*}
E_{n}(\alpha, \lambda)=-(\lambda-n)^{2}+\alpha\left(\lambda+n^{2}\right)=-\left[\frac{1}{2}(\Delta-1)-n\right]^{2}+\alpha n(n+1)+\frac{1}{4} \alpha^{2} \tag{4.14}
\end{equation*}
$$

and
$\psi_{n}(\alpha, \lambda ; x)=N_{n}(\alpha, \lambda)(\operatorname{sech} x)^{\frac{\lambda}{1-\alpha}-n}\left(1+\alpha \sinh ^{2} x\right)^{\frac{1}{2}\left(\frac{\lambda}{1-\alpha}-n-1\right)} P_{n}(\alpha, \lambda ; \tanh x)$
where $\lambda$ is defined by $\lambda=\frac{1}{2}(\alpha-1+\Delta), \Delta \equiv \sqrt{(1-\alpha)^{2}+4 A(A+1)}$.
For any $n=0,1,2, \ldots$, the function (4.15) is square integrable because $\psi_{n}(\alpha, \lambda ; x) \sim$ $\mathrm{e}^{-|x|}$ for $x \rightarrow \pm \infty$. However, in the same limits, the deforming function $f(\alpha ; x)$, given by (2.2) and (4.12), behaves as $\mathrm{e}^{2|x|}$. Hence condition (2.27), necessary to ensure the Hermiticity of $\pi$, cannot be satisfied. From this we infer that with a deforming function corresponding to (4.12), the hyperbolic Pöschl-Teller potential has no bound state. The same result remains valid for the free-particle problem and contrasts with what was obtained in [19] in another
context. While this shows that the result is strictly environment dependent, one must also remember that in the conventional free-particle problem (see, e.g., [43]) one way to avoid the divergence is to assume that the particle is confined to a closed and finite universe. In the context of PDEM a similar philosophy may be adopted with regard to the preservation of condition (2.27).

In conclusion, we have shown how the simple fact of going from trigonometric to hyperbolic functions in a deformed or PDEM environment may drastically change the picture as far as an ES potential bound-state spectrum is concerned. In this respect, the new condition (2.27) or (2.28), introduced in this paper, has played an essential role.

## 5. Results for shape-invariant potentials

The procedure demonstrated on some simple examples in section 4 can be easily generalized to other shape-invariant potentials. In the appendix, we list some of the results obtained when taking for $V_{\text {eff }}(\boldsymbol{b} ; x)$ the potentials considered in table 4.1 of [24]. In this respect, two important remarks are in order.

First, although the trigonometric Pöschl-Teller potential of section 4.1 may be considered as a limiting case of Rosen-Morse I potential when its parameter $B$ goes to zero (and a change of variable $x \in[0, \pi] \rightarrow x^{\prime}=x-\frac{\pi}{2} \in\left[-\frac{\pi}{2}, \frac{\pi}{2}\right]$ is performed), the different choices of deforming function made in section 4.1 and in the appendix produce unrelated results in the deformed or PDEM context.

Second, three of the potentials listed in [24] are missing from the appendix, namely Scarf II, Rosen-Morse II and generalized Pöschl-Teller potentials. The reasons for their absence are different. For Scarf II potential, it turns out that no nontrivial values of the parameters may ensure positive definiteness of $f(\boldsymbol{\alpha} ; x)$ on the whole real line, on which the potential is defined. On the other hand, for Rosen-Morse II and generalized Pöschl-Teller potentials the resulting square-integrable wavefunctions do not ensure the Hermiticity of $\pi$, as expressed in condition (2.27). As a consequence, these potentials, which both have a finite number of bound states in the undeformed case, do not support any bound state in the deformed one.

Let us now turn ourselves to the results listed in the appendix. For the SI potentials $V_{\text {eff }}(\boldsymbol{b} ; x)$ considered there, the potentials $V(\boldsymbol{a} ; x)$ to be used in the PDEM equation (2.3) fall into two categories. For the shifted oscillator, three-dimensional oscillator, Coulomb and Morse potentials, $V(\boldsymbol{a} ; x)$ has the same shape as $V_{\text {eff }}(\boldsymbol{b} ; x)$. The only effect of the mass and ambiguity parameters indeed amounts to a renormalization of the potential parameters and/or an energy shift $\delta v$. So we obtain

$$
\begin{align*}
& V(\boldsymbol{a} ; x)=\frac{1}{4} \omega^{* 2}\left(x-\frac{2 b^{*}}{\omega^{*}}\right)^{2}+\delta v  \tag{5.1}\\
& V(\boldsymbol{a} ; x)=\frac{1}{4} \omega^{* 2} x^{2}+\frac{l(l+1)}{x^{2}}+\delta v  \tag{5.2}\\
& V(\boldsymbol{a} ; x)=-\frac{\mathrm{e}^{2}}{x}+\frac{l(l+1)}{x^{2}}+\delta v  \tag{5.3}\\
& V(\boldsymbol{a} ; x)=B^{* 2} \mathrm{e}^{-2 x}-B^{*}\left(2 A^{*}+1\right) \mathrm{e}^{-x} \tag{5.4}
\end{align*}
$$

respectively, where, for Morse potential, for instance,

$$
\begin{equation*}
A^{*}=\frac{1}{2}\left(\frac{B(2 A+1)+\rho \alpha}{\sqrt{B^{2}-(\rho+\sigma) \alpha^{2}}}-1\right) \quad B^{*}=\sqrt{B^{2}-(\rho+\sigma) \alpha^{2}} \tag{5.5}
\end{equation*}
$$

One may observe strikingly distinct influences of deformation or mass parameters on bound-state energy spectra. In some cases (shifted oscillator, three-dimensional oscillator, Scarf I and Rosen-Morse I), the infinite number of bound states of conventional quantum mechanics remains infinite after the onset of deformation. Similarly, for Morse potential and for Eckart potential with $\alpha \neq-2$, one keeps a finite number of bound states. For the Coulomb potential, however, the infinite number of bound states is converted into a finite one, while for Eckart potential with $\alpha=-2$, the finite number of bound states becomes infinite. It is also remarkable that, whenever finite, the bound-state number becomes dependent on the deforming parameter.

In the appendix, for lack of space we have not shown the explicit form of the excited-state wavefunctions, in particular that of the polynomials $P_{n}(\boldsymbol{\alpha}, \boldsymbol{\lambda} ; y)^{4}$. For the same reason, we have not exhibited the SUSY partners of $V_{\text {eff }}(\boldsymbol{b} ; x)$, which can be easily determined from equation (2.20) and reduce to the conventional ones of [24] in the constant-mass limit.

## 6. Conclusion

In this paper, we have generated a lot of new ES potentials associated with a PDEM background. For such a purpose, we have considered known SI potentials for the constant-mass SE as effective potentials in the PDEM one, taking into account the ambiguity-parameterdependent contribution coming from the momentum and mass-operator noncommutativity. The corresponding deformed SI condition solvability has imposed the general form of both the deformed superpotential and the PDEM. For the latter, we have then chosen a fairly general particular case and we have found both the corresponding ES potential and the bound-state energy spectrum and wavefunctions.

The existence of such a spectrum is determined not only by a square-integrability condition on the wavefunctions as in conventional quantum mechanics, but also by a Hermiticity condition on the deformed momentum operator. The latter is a new and important contribution of this paper. As we have demonstrated on some specific examples, it may have relevant effects whenever the PDEM vanishes at an end point of the interval on which the potential is defined.

We have shown that in some cases the new ES potential has the same shape as the conventional SI potential used in the construction, but that in others the ambiguity-parameterdependent term turns out to change its shape.

Furthermore, if in some instances the spectrum of the new ES potential results from a smooth deformation of that of the conventional SI one, we have also observed in other examples a generation or suppression of bound states, depending on the values taken by the mass parameters. We would like to stress the nontrivial nature of this result, of which very few cases have been signalled in the literature devoted to PDEM SEs so far.

It is rather obvious that our results for bound states could be easily extended to the $S$ matrix and that our construction method of new ES PDEM potentials could also be applied to more complicated forms of the PDEMs or to other potentials that are SI under parameter translation. An interesting open question for future work is whether it could be generalized to other types of SI, such as SI under parameter scaling.

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[^0]
## Appendix

In this appendix, we list some of the results obtained for the SI potentials considered in table 4.1 of [24] when deforming the corresponding SI condition as explained in sections 2 and 3. For simplicity's sake, the parameter dependence of the functions has not been indicated explicitly.
Shifted oscillator
$V_{\text {eff }}(x)=\frac{1}{4} \omega^{2}\left(x-\frac{2 b}{\omega}\right)^{2} \quad W(x)=\lambda x+\mu \quad($ class 1: $\phi(x)=x)$
$g(x)=\alpha x^{2}+2 \beta x \quad \alpha>\beta^{2} \geqslant 0 \quad \lambda=\frac{1}{2}(\alpha+\Delta) \quad \mu=\beta-\frac{b \omega}{2 \lambda}$
$\Delta \equiv \sqrt{\omega^{2}+\alpha^{2}} \quad \lambda_{i}=\lambda+i \alpha \quad \mu_{i}=\frac{\lambda \mu+2 i \beta \lambda+i^{2} \alpha \beta}{\lambda+i \alpha}$
$E_{n}=\left(n+\frac{1}{2}\right) \Delta+\left(n^{2}+n+\frac{1}{2}\right) \alpha+b^{2}-\left(\frac{\left[(2 n+1) \Delta+\left(2 n^{2}+2 n+1\right) \alpha\right] \beta-b \omega}{\Delta+(2 n+1) \alpha}\right)^{2}$
$n=0,1,2, \ldots$
$\psi_{0}(x) \propto f^{-(\lambda+\alpha) /(2 \alpha)} \exp \left(\frac{\lambda \beta-\mu \alpha}{\alpha \delta} \arctan \frac{\alpha x+\beta}{\delta}\right) \quad \delta \equiv \sqrt{\alpha-\beta^{2}}$
$\tilde{V}(x)=2(\rho+2 \sigma) \alpha x(\alpha x+2 \beta)+2 \rho \alpha+4 \sigma \beta^{2}$.
Three-dimensional oscillator
$V_{\text {eff }}(x)=\frac{1}{4} \omega^{2} x^{2}+\frac{l(l+1)}{x^{2}} \quad 0 \leqslant x<\infty$
$W(x)=\frac{\lambda}{x}+\mu x \quad\left(\right.$ class 2: $\left.\phi(x)=\frac{1}{x}\right) \quad g(x)=\alpha x^{2} \quad \alpha>0$
$\lambda=-l-1 \quad \mu=\frac{1}{2}(\alpha+\Delta) \quad \Delta \equiv \sqrt{\omega^{2}+\alpha^{2}}$
$\lambda_{i}=\lambda-i \quad \mu_{i}=\mu+i \alpha$
$E_{n}=\Delta\left(2 n+l+\frac{3}{2}\right)+\alpha\left[2(n+l+1)(2 n+1)+\frac{1}{2}\right] \quad n=0,1,2, \ldots$
$\psi_{0}(x) \propto x^{l+1} f^{-[\mu+(l+2) \alpha] /(2 \alpha)} \quad \tilde{V}(x)=2(\rho+2 \sigma) \alpha^{2} x^{2}+2 \rho \alpha$.

## Coulomb

$V_{\text {eff }}(x)=-\frac{e^{2}}{x}+\frac{l(l+1)}{x^{2}} \quad 0 \leqslant x<\infty \quad W(x)=\frac{\lambda}{x}+\mu \quad\left(\right.$ class $\left.1: \phi(x)=\frac{1}{x}\right)$
$g(x)=\alpha x \quad \alpha>0 \quad \lambda=-l-1 \quad \mu=-\frac{\mathrm{e}^{2}+\alpha \lambda}{2 \lambda}$
$\lambda_{i}=\lambda-i \quad \mu_{i}=-\frac{\mathrm{e}^{2}+\alpha \lambda(2 i+1)-\alpha i^{2}}{2(\lambda-i)}$
$E_{n}=-\left(\frac{\mathrm{e}^{2}-\alpha\left[n^{2}+(l+1)(2 n+1)\right]}{2(n+l+1)}\right)^{2}$
$n=0,1, \ldots, n_{\max }$, where $n_{\max }=$ largest integer such that
$n^{2}+(l+1)(2 n+1)<\frac{\mathrm{e}^{2}}{\alpha}$ if $\alpha<\frac{\mathrm{e}^{2}}{l+1}$
$\psi_{0}(x) \propto x^{l+1} f^{-\left(\frac{\mu}{\alpha}+l+\frac{3}{2}\right)} \quad \tilde{V}(x)=\sigma \alpha^{2}$.

## Morse

$V_{\text {eff }}(x)=B^{2} \mathrm{e}^{-2 x}-B(2 A+1) \mathrm{e}^{-x} \quad A, B>0$
$W(x)=\lambda \mathrm{e}^{-x}+\mu \quad\left(\right.$ class 1: $\left.\phi(x)=\mathrm{e}^{-x}\right) \quad g(x)=\alpha \mathrm{e}^{-x} \quad \alpha>0$
$\lambda=-\frac{1}{2}(\alpha+\Delta) \quad \mu=-\frac{1}{2}\left(\frac{B(2 A+1)}{\lambda}+1\right) \quad \Delta \equiv \sqrt{4 B^{2}+\alpha^{2}}$
$\lambda_{i}=\lambda-i \alpha \quad \mu_{i}=\frac{2 \lambda(\mu-i)+i^{2} \alpha}{2(\lambda-i \alpha)}$.
$E_{n}=-\frac{1}{4}\left(\frac{2 B(2 A+1)-\left[(2 n+1) \Delta+\left(2 n^{2}+2 n+1\right) \alpha\right]}{\Delta+(2 n+1) \alpha}\right)^{2}$
$n=0,1, \ldots, n_{\text {max }}$, where $n_{\text {max }}=$ largest integer smaller than $A$
and such that $\alpha<\alpha_{\max }\left(n_{\max }\right)$ with
$\alpha_{\text {max }}(0)=\frac{4 A(A+1) B}{2 A+1}$
$\alpha_{\max }(n)=\frac{B(2 A+1)\left(2 n^{2}+2 n+1\right)-B(2 n+1)\left[(2 A+1)^{2}+4 n^{2}(n+1)^{2}\right]^{1 / 2}}{2 n^{2}(n+1)^{2}}$
$n=1,2, \ldots$
$\psi_{0}(x) \propto f^{\frac{\lambda}{\alpha}-\mu-\frac{1}{2}} \mathrm{e}^{-\mu x}$
$\tilde{V}(x)=(\rho+\sigma) \alpha^{2} \mathrm{e}^{-2 x}+\rho \alpha \mathrm{e}^{-x}$

## Eckart

$V_{\text {eff }}(x)=A(A-1) \operatorname{csch}^{2} x-2 B \operatorname{coth} x \quad A \geqslant \frac{3}{2} \quad B>A^{2} \quad 0 \leqslant x<\infty$
$W(x)=\lambda \operatorname{coth} x+\mu \quad($ class $1: \phi(x)=\operatorname{coth} x)$
$g(x)=\alpha \mathrm{e}^{-x} \sinh x \quad-2 \leqslant \alpha \neq 0 \quad \lambda=-A \quad \mu=\frac{B}{A}-\frac{1}{2} \alpha$
$\lambda_{i}=\lambda-i \quad \mu_{i}=\frac{\lambda \mu-\frac{1}{2} \alpha i(2 \lambda-i)}{\lambda-i}$.
$E_{n}=-(A+n)^{2}-\left(\frac{B-\frac{1}{2} \alpha\left[(2 n+1) A+n^{2}\right]}{A+n}\right)^{2}-\alpha\left[(2 n+1) A+n^{2}\right]$
$n=0,1,2, \ldots$ if $\alpha=-2$
$n=0,1, \ldots, n_{\max }$ if $\alpha>-2$, where $n_{\max }=$ largest integer such that
$(A+n)^{2}<\frac{2 B+\alpha A(A-1)}{2+\alpha}$
$\psi_{0}(x) \propto(\operatorname{coth} x-1)^{-A-1} \operatorname{csch} x \exp \left(-\frac{\mu-A}{\operatorname{coth} x-1}\right) \quad$ if $\alpha=-2$
$\propto(\operatorname{coth} x+1)^{1 / 2}(\operatorname{coth} x+1+\alpha)^{-\frac{(1+\alpha)+\mu+\mu}{2+\alpha}-\frac{1}{2}}(\operatorname{coth} x-1)^{\frac{\mu-A}{2+\alpha}} \quad$ if $\alpha>-2$
$\tilde{V}(x)=(\rho+\sigma) \alpha^{2} \mathrm{e}^{-4 x}-\rho \alpha(2+\alpha) \mathrm{e}^{-2 x}$.

## Scarf I

$V_{\text {eff }}(x)=\left(B^{2}+A^{2}-A\right) \sec ^{2} x-B(2 A-1) \tan x \sec x \quad 0<B<A-1 \quad-\frac{\pi}{2} \leqslant x \leqslant \frac{\pi}{2}$
$W(x)=\lambda \tan x+\mu \sec x \quad($ class 3: $\phi(x)=\sin x) \quad g(x)=\alpha \sin x \quad 0<|\alpha|<1$
$\lambda=\frac{1}{2}\left(1+\Delta_{+}+\Delta_{-}\right) \quad \mu=\frac{1}{2}\left(\alpha-\Delta_{+}+\Delta_{-}\right)$
$\Delta_{ \pm} \equiv \sqrt{\frac{1}{4}(1 \mp \alpha)^{2}+(A \pm B)(A \pm B-1)} \quad \lambda_{i}=\lambda+i \quad \mu_{i}=\mu+i \alpha$
$E_{n}=-\frac{1}{4}\left(2 n+1+\Delta_{+}+\Delta_{-}\right)^{2}+\alpha\left(n+\frac{1}{2}\right)\left(\Delta_{+}-\Delta_{-}\right)-\alpha^{2}\left(n^{2}+n+\frac{1}{2}\right) \quad n=0,1,2, \ldots$
$\psi_{0}(x) \propto f^{-\frac{\lambda-\alpha \mu}{1-\alpha^{2}}-\frac{1}{2}}(1-\sin x)^{\frac{\lambda+\mu}{2(1+\alpha)}}(1+\sin x)^{\frac{\lambda-\mu}{2(1-\alpha)}}$
$\tilde{V}(x)=-(\rho+\sigma) \alpha^{2} \sin ^{2} x-\rho \alpha \sin x+\sigma \alpha^{2}$.

## Rosen-Morse I

$V_{\text {eff }}(x)=A(A-1) \csc ^{2} x+2 B \cot x \quad A \geqslant \frac{3}{2} \quad 0 \leqslant x \leqslant \pi$
$W(x)=\lambda \cot x+\mu \quad($ class 1: $\phi(x)=\cot x)$
$g(x)=\sin x(\alpha \cos x+\beta \sin x) \quad \frac{|\alpha|}{2}<\sqrt{1+\beta} \quad \beta>-1$
$\lambda=-A \quad \mu=-\frac{B}{A}-\frac{1}{2} \alpha \quad \lambda_{i}=\lambda-i \quad \mu_{i}=\frac{\lambda \mu-\frac{1}{2} \alpha i(2 \lambda-i)}{\lambda-i}$
$E_{n}=(A+n)^{2}-\left(\frac{B+\frac{1}{2} \alpha\left[(2 n+1) A+n^{2}\right]}{A+n}\right)^{2}+\beta\left[(2 n+1) A+n^{2}\right] \quad n=0,1,2, \ldots$
$\psi_{0}(x) \propto f^{-(A+1) / 2}(\sin x)^{A} \exp \left(\frac{\mu+\frac{1}{2} \alpha A}{\delta} \arctan \frac{\cot x+\frac{\alpha}{2}}{\delta}\right) \quad \delta \equiv \sqrt{1+\beta-\frac{\alpha^{2}}{4}}$
$\tilde{V}(x)=(\rho+\sigma)\left[\frac{1}{2}\left(\alpha^{2}-\beta^{2}\right) \cos 4 x+\alpha \beta \sin 4 x\right]+\rho(2+\beta)(-\alpha \sin 2 x+\beta \cos 2 x)$
$+(-\rho+\sigma) \frac{1}{2}\left(\alpha^{2}+\beta^{2}\right)$.

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[^0]:    ${ }^{4}$ Detailed results are available from the authors.

